## Stellar cosmic ray driven chemistry in the terrestrial planet-forming region of protoplanetary disks

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### **Ionisation sources in protoplanetary disks**

• Stellar cosmic rays:

Turner & Drake (2009); Rab et al (2017); Rodgers-Lee et al (2017, 2020a); Fraschetti et al (2018), Offner et al. (2019)

- Galactic cosmic rays
- Stellar X-ray & FUV photons
- Radioactive nuclides (Brandt Gaches' talk!)



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# Galactic cosmic rays are suppressed by magnetised stellar winds



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## Can stellar cosmic rays explain low CO abundances observed by ALMA?



- Low CO abundances at large radii in protoplanetary disks (Miotello et al 2017)
- Schwarz et al (2018): Chemical processing not enough

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Alexandra Mulholland (undergrad. research project)

# Faster diffusion leads to higher ionisation rate at larger distances





 $\rightarrow$  Considering GeV energy cosmic rays and diffusive transport

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# Stellar cosmic rays deplete CO in the inner terrestrial planet-forming region





# Stellar cosmic rays deplete CO in the inner terrestrial planet-forming region





# Stellar cosmic rays deplete CO in the inner terrestrial planet-forming region





## CO is significantly depleted within 1 Myr



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### Stellar cosmic ray effect on most abundant species ISM GCRs



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### Stellar cosmic ray effect on most abundant species ISM GCRs



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## Stellar cosmic ray ionisation rates agree better with Spitzer observations

- Spitzer fluxes (Banzatti et al 2020)
- Synthetic spectra using Slabspec (Salyk et al 2020)
- 15 micron HCN band and 14 micron CO<sub>2</sub> band



### **Next steps**



### Can we explain radially varying ionisation rates? (Seifert et al 2021)

#### What is the radial dependence of the Galactic cosmic ray spectrum?

#### What about transition disks?

## Conclusions



GeV energy stellar cosmic rays can significantly deplete CO in the terrestrial planet-forming region of protoplanetary disks within 3 Myr

Spitzer flux ratio measurements of HCN and CO<sub>2</sub> are better matched when stellar cosmic rays are included

JWST's MIRI may contribute to our understanding of stellar cosmic ray fluxes

Go raibh maith agaibh (Thank you!) Any questions → dlee@cp.dias.ie