## GALAXY EVOLUTION: MODELING THE ROLE OF NON-THERMAL PRESSURE IN THE INTERSTELLAR MEDIUM

BIRNBOIM, BALBERG & TEYSSIER MNRAS 2015 447, 3678

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## MAIN POINTS:

1. FOR ISM, NON-THERMAL PRESSURE IS

"AT LEAST AS IMPORTANT" AS THERMAL

N.T. ENERGY: MAGNETIC, COSMIC RAYS,

TURBULENCE

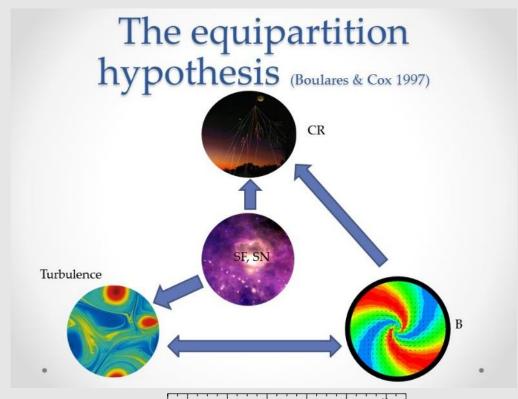
EFFECT: KEEP THE ISM DIFFUSE

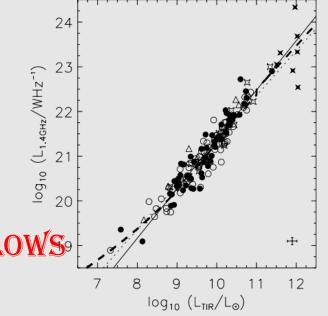
- 2. RADIO-IR RELATION IS USED TO CALIBRATE N.T. ENERGY SFR RELATION
- 3. COARSE GRAINED N.T. CAN BE MODELED AS SINGLE N.T. PRESSURE.

SUCCESS! WEAKER FEEDBACK,

CORRECT GAS DEPLETION RATES,

SELF REGULATION WITH WEAKER OUTFLOWS9





## WHATS IN THE PAPER?

QUASI ANALYTIC MODEL

RAMSES 3D WITH DEDICATED SUBGRID MODEL

PRELIMINARY RESULTS

