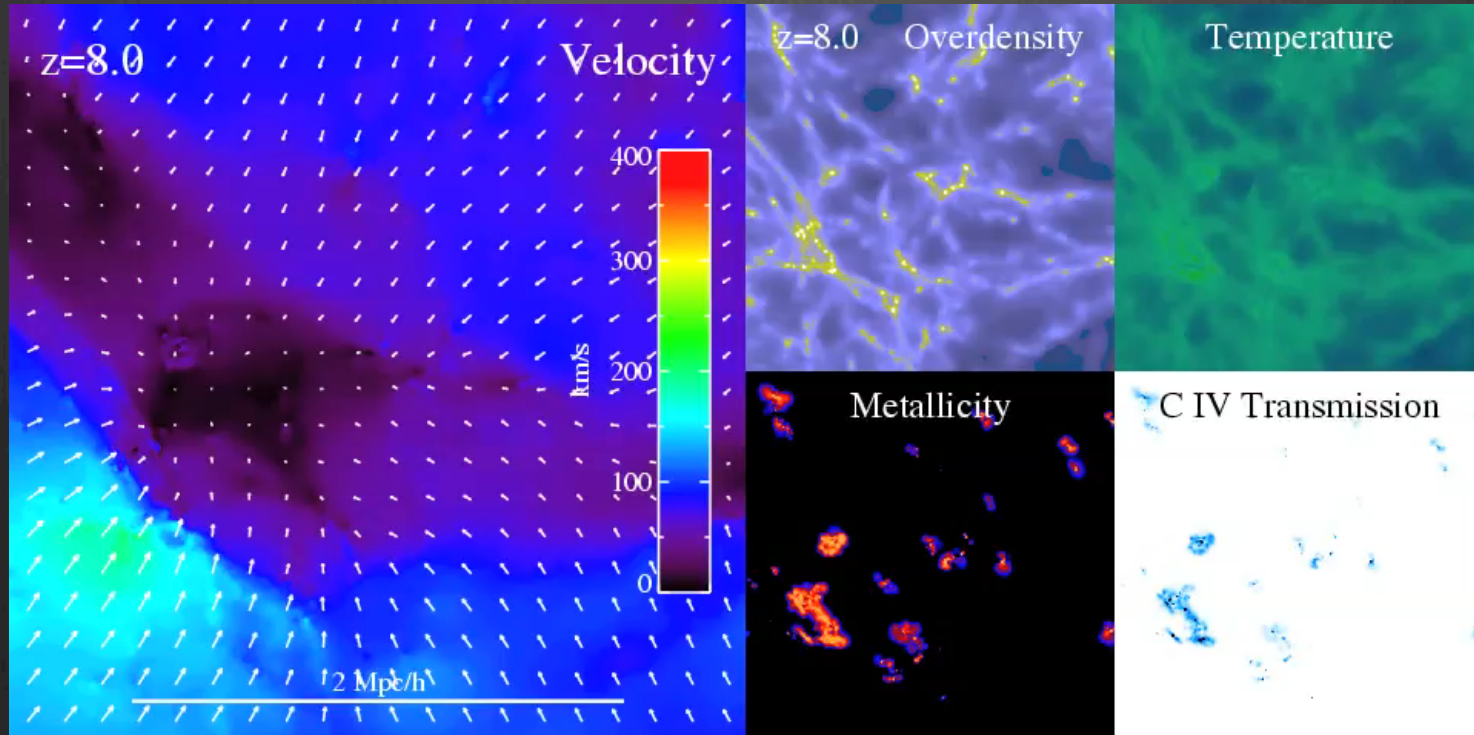


Constraining Inflows and Outflows From Galaxy Growth Across Cosmic Time



Romeel Davé, UWC/SAAO/AIMS (Cape Town)

with Sourav Mitra & Kristian Finlator

and R Somerville, B Oppenheimer, R Thompson



Includes:

Gravity
Gas (moving mesh)
Star formation
Photoionizing bkgd
Black hole growth
10-species chemistry
Type II SNe feedback
Type Ia SNe feedback
AGB stellar evol
AGN feedback



Stars Gas density
Temp. Metallicity

~6 billion cells
1 kpc resolution
100 Mpc volume

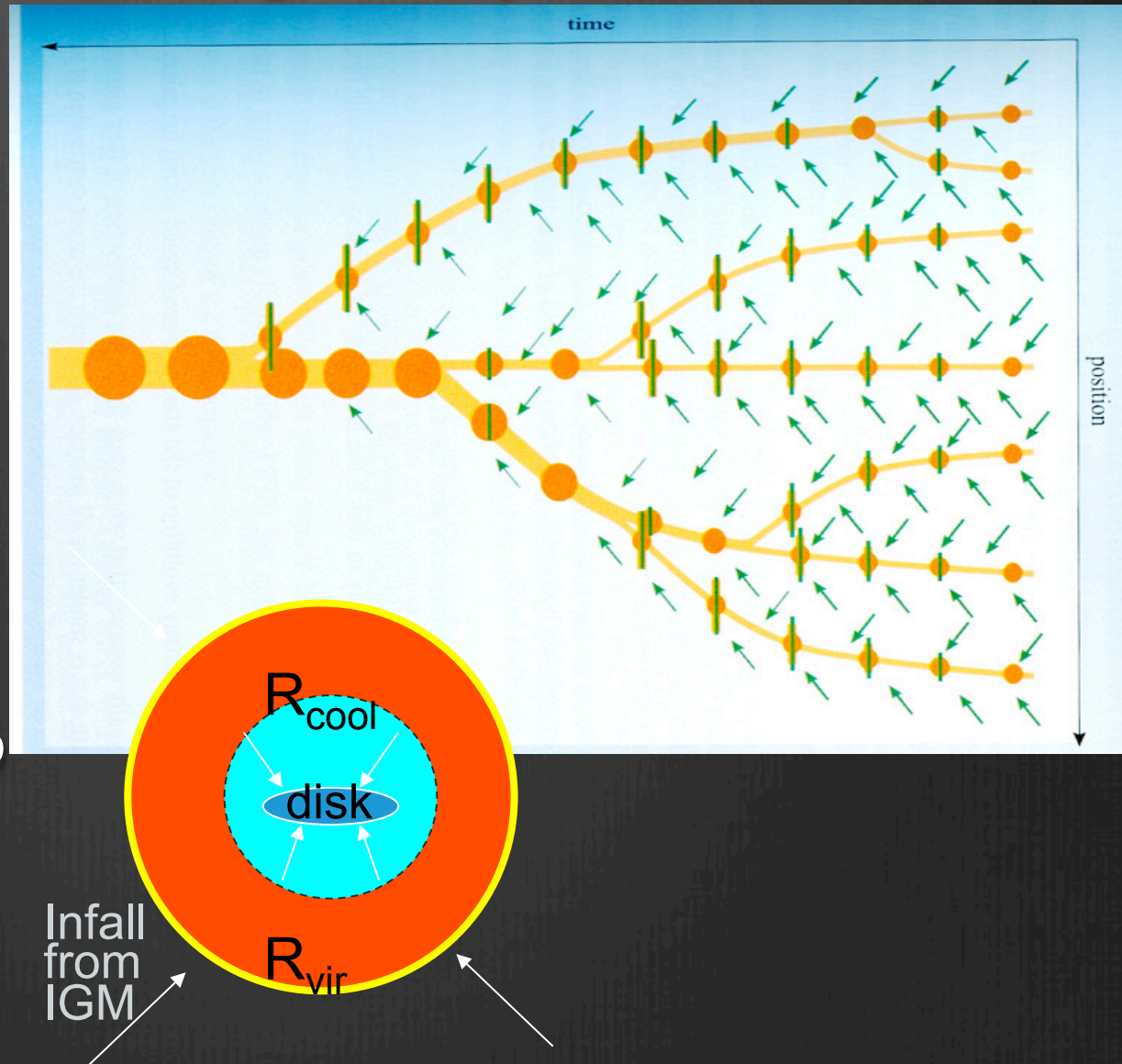
The Illustris Project: AREPO

SAM cartoon of galaxy formation

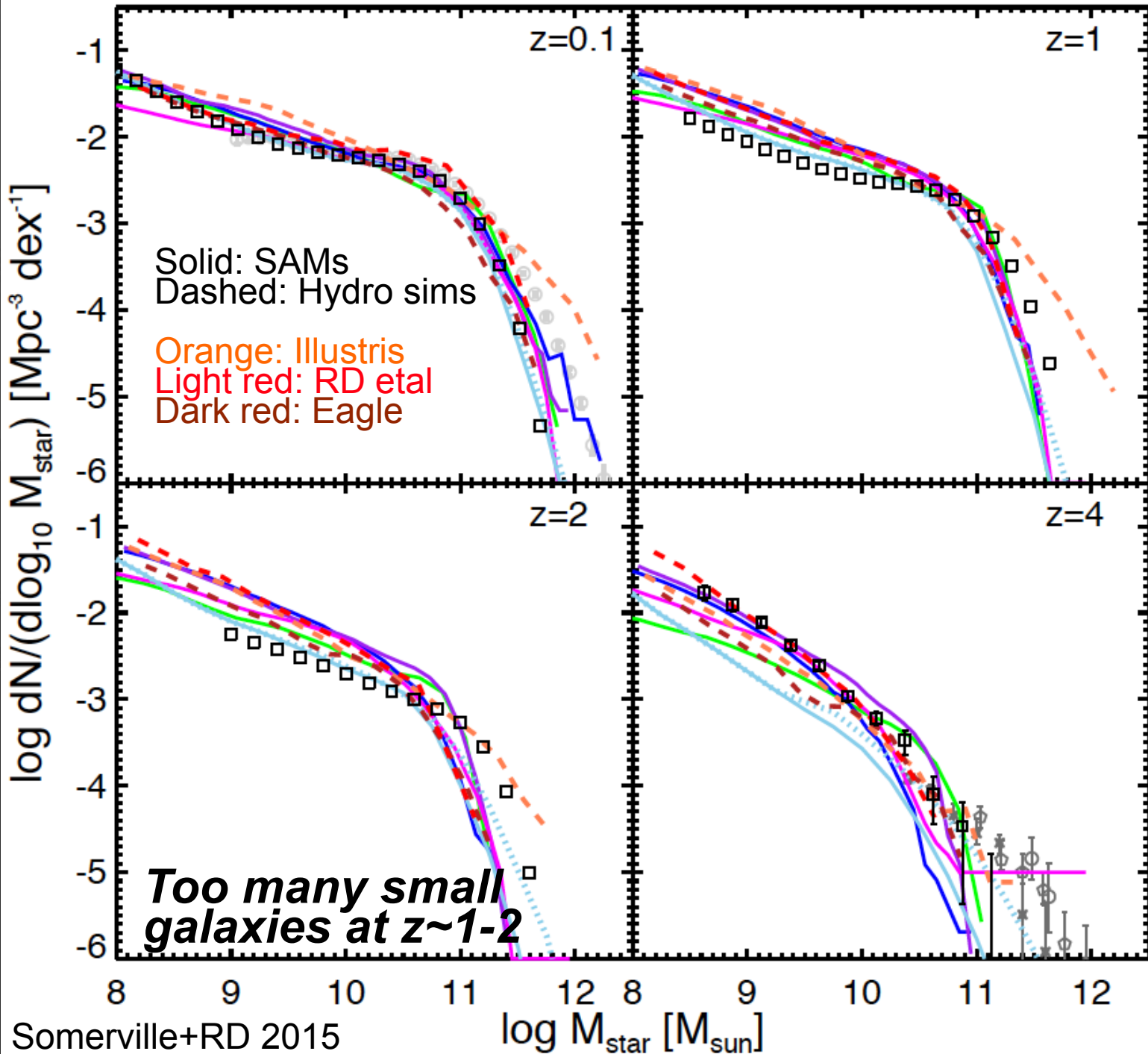
Take N-body merger tree, add parameterized baryonic physics:

SFR, cooling, feedback, galaxy merging, ...

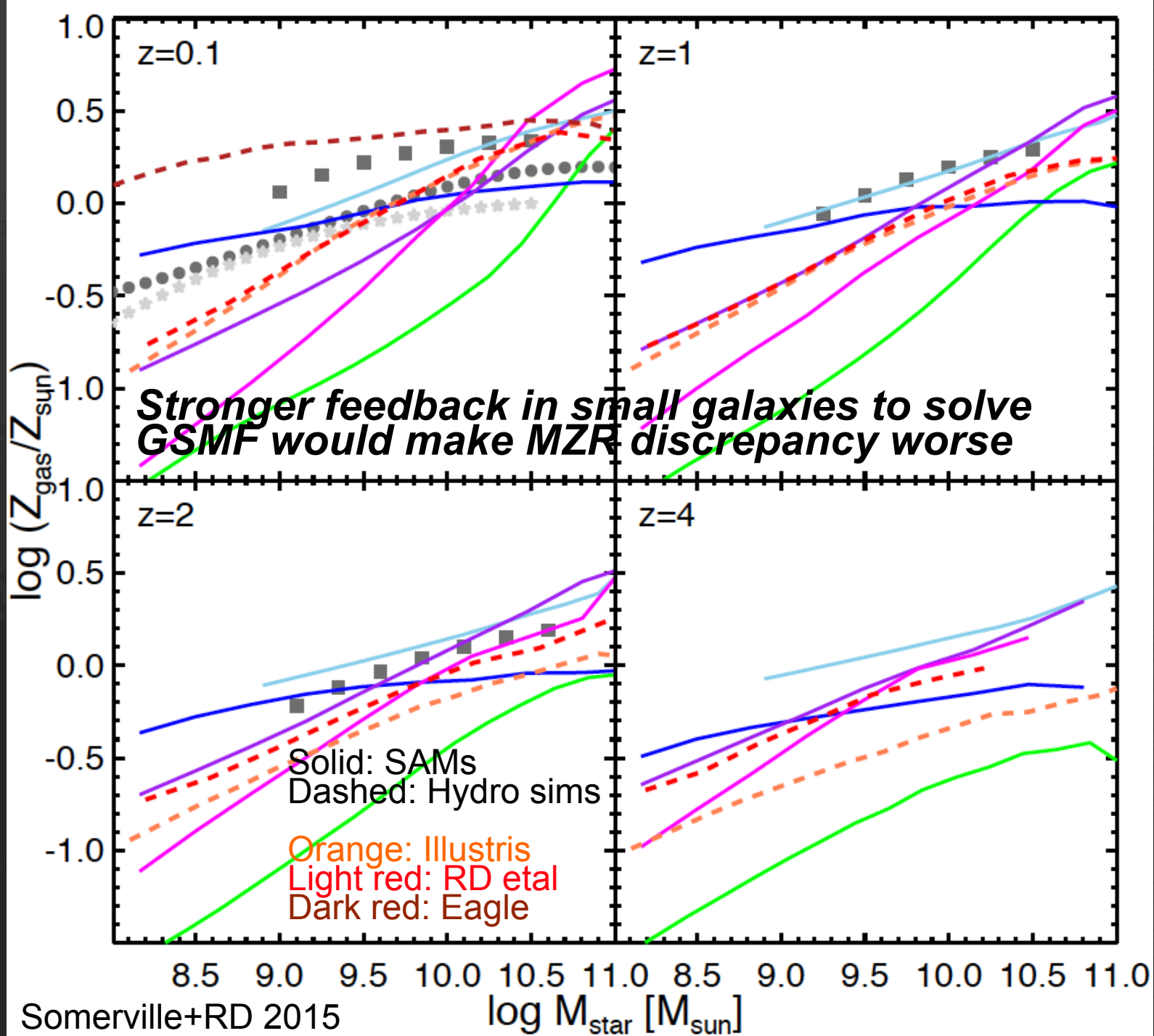
Typically ~ 40 - 50 parameters, tuned to match $z=0$ data.



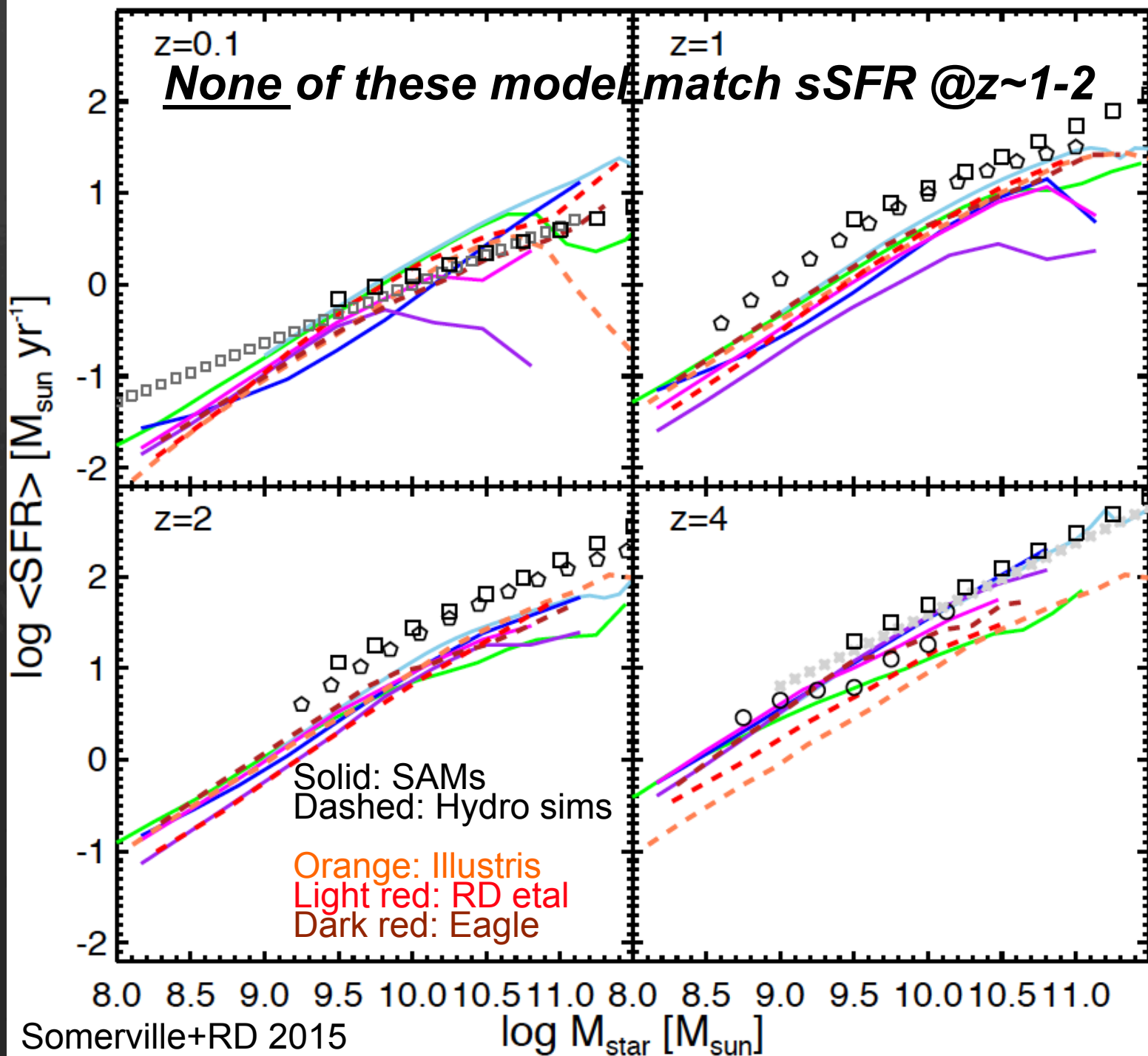
Galaxy Stellar Mass Function

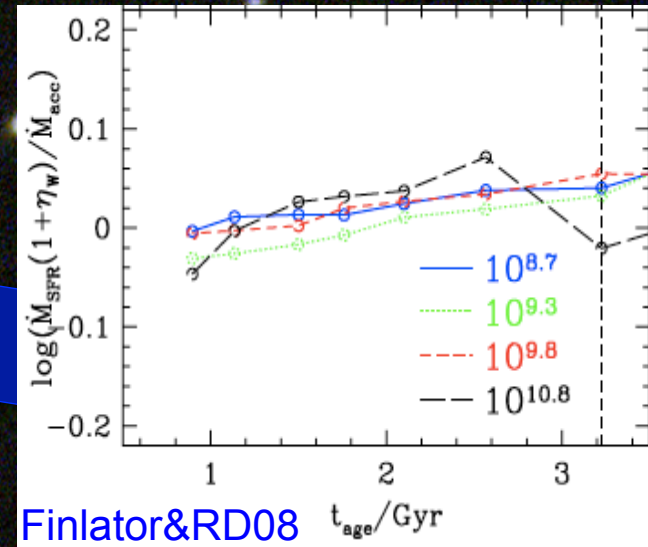
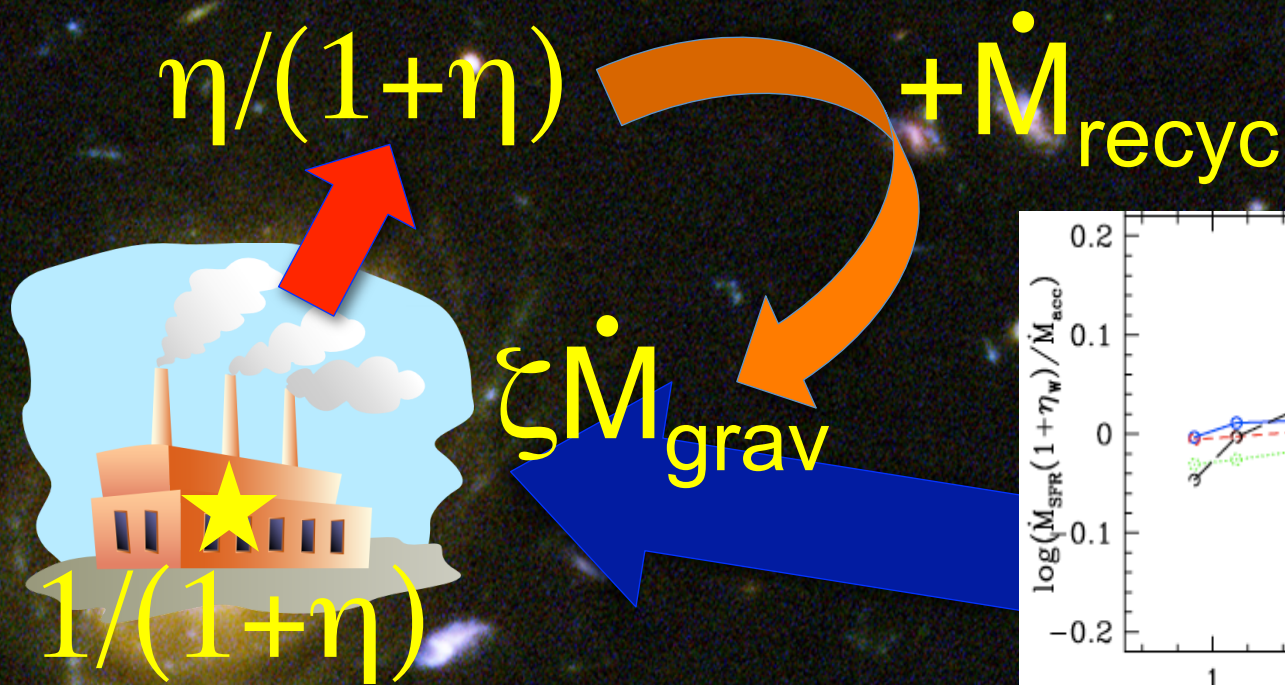


Mass-Metallicity Relation



Galaxy "Main Sequence"





(FROM TGM)

Equilibrium Model:

Inflow = SFR + Outflow + $d\text{Reservoir}/dt$

$$\text{SFR} = (\zeta\dot{M}_{\text{grav}} + \dot{M}_{\text{recyc}})/(1+\eta)$$

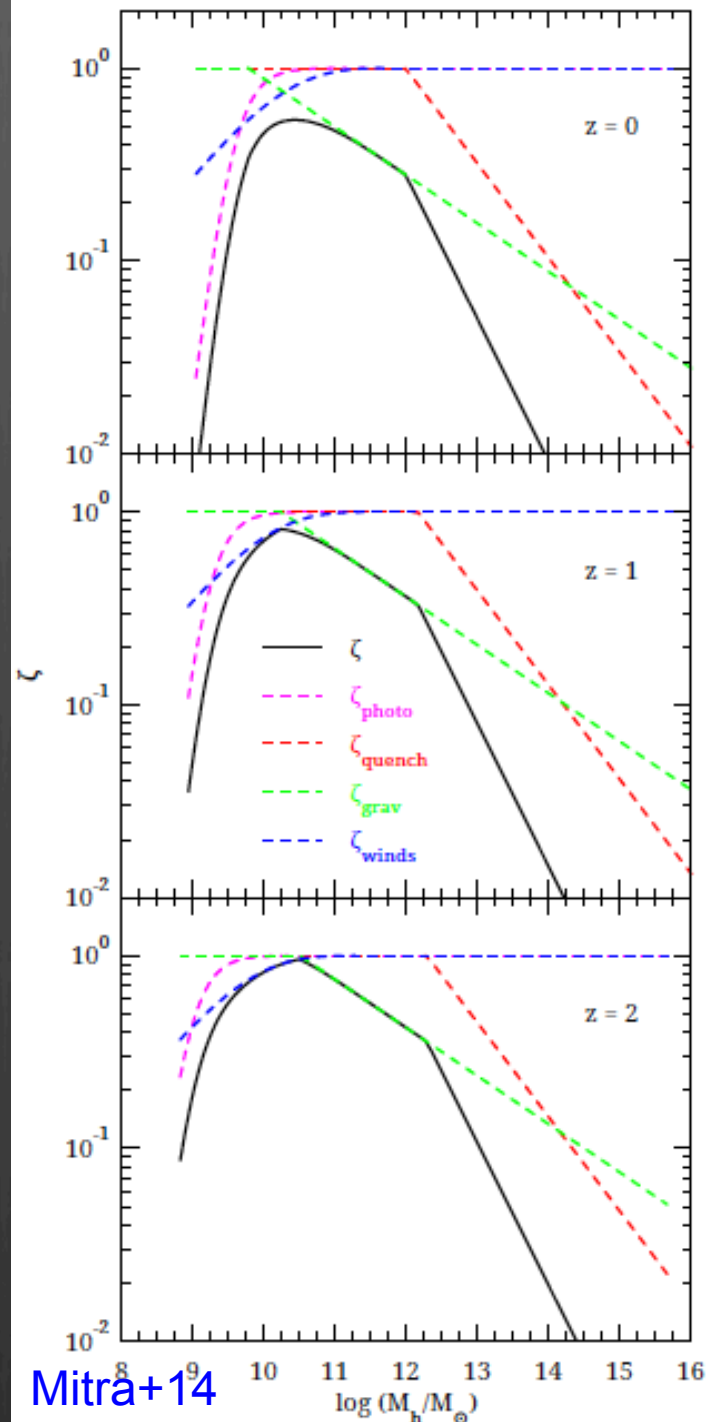
$$Z = y \text{SFR} / \zeta\dot{M}_{\text{grav}}$$

Finlator+08
 Bouché+10
 Dave+11,12
 Lilly+13
 Dekel+14

Baryon cycling parameters

- *Ejective feedback*
 $\eta = \text{Outflow/SFR}$
- *Preventive feedback*
 ζ (J_{UV} , AGN, ...)
- *Wind recycling time*
 $t_{\text{rec}} = \text{time for outflow to return.}$

Each depends on
 $M_{\text{halo}}, z, \dots?$



Parameterize, Bayesian MCMC

- ⊛ Turns out we need 8 parameters (Bayesian evidence analysis shows that removing any more is not preferred):

$$\eta = \left(\frac{M_h}{10^{\eta_1 + \eta_2 \sqrt{z}}} \right)^{\eta_3}$$

Mitra+14

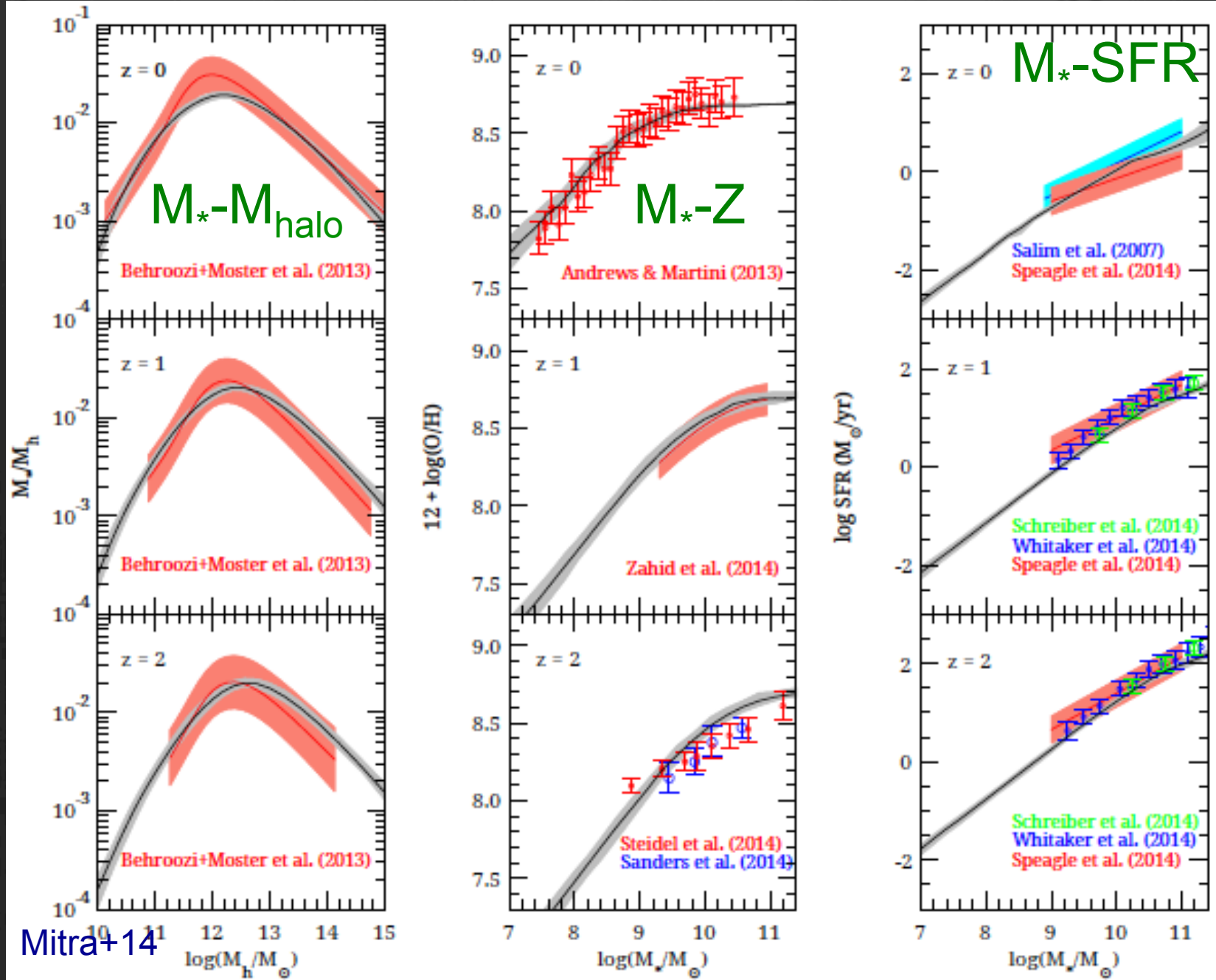
$$\zeta_{\text{quench}} = \text{MIN} \left[1, \left(\frac{M_h}{M_q} \right)^{\zeta_1} \right], \frac{M_q}{10^{12} M_{\odot}} = (0.96 + \zeta_2 z).$$

$$t_{\text{rec}} = \tau_1 \times 10^9 \text{ yr} \times (1 + z)^{\tau_2} \left(\frac{M_h}{10^{12}} \right)^{\tau_3}$$

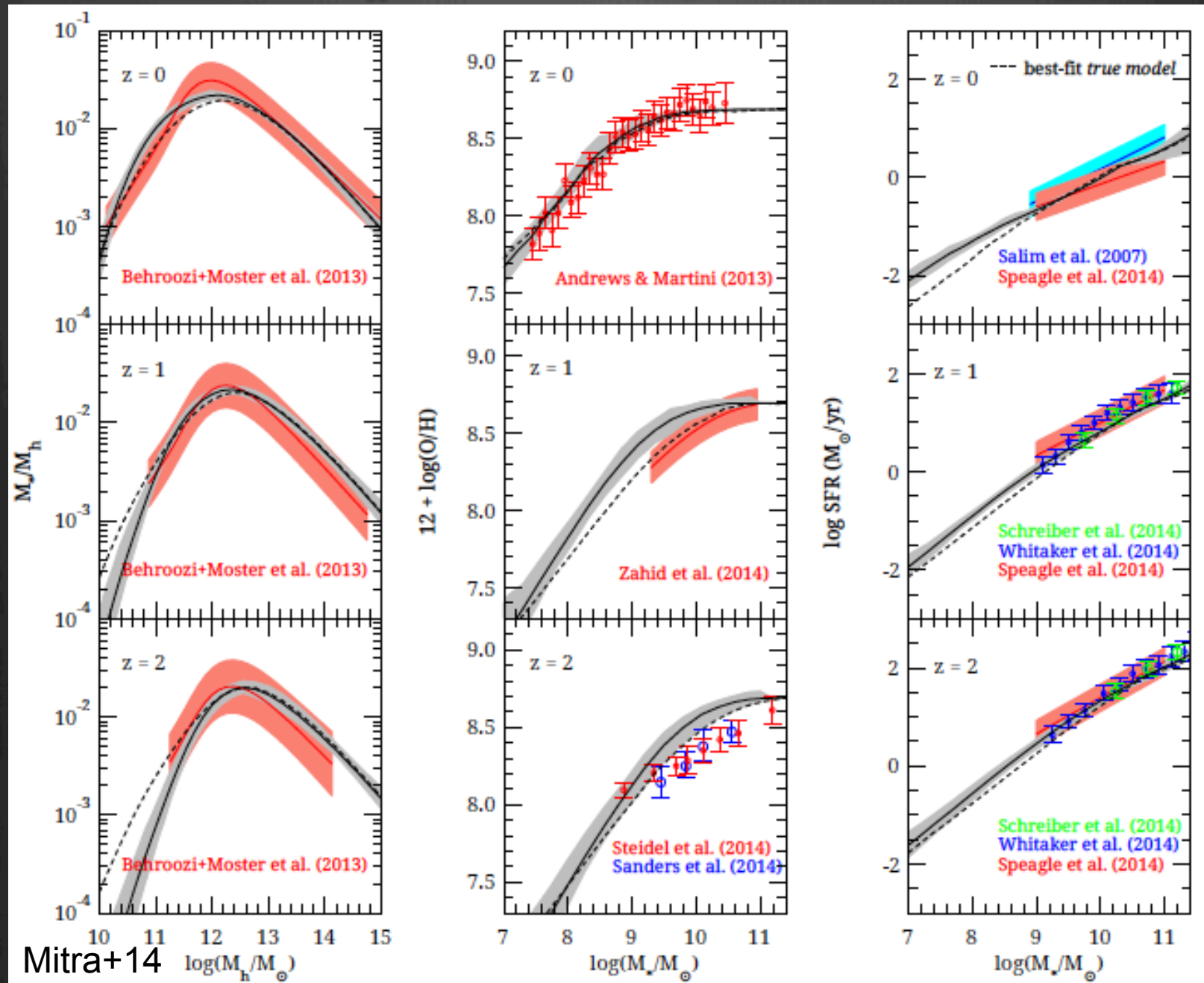
Observational Constraints

- **M_* - M_{halo}** (equivalent to GSMF, $z=0-2$): Inferred from GSMF($z\sim 0-6$) data, consistent w/SFR evol. [Behroozi+13, Moster+13].
- **Mass-Metallicity relation:** Now seen out to $z\sim 2$ thanks to Keck/Mosfire. [Steidel+14, Sanders+14]
- **SFR- M_*** ($z=0-2$): Recent compilation with consistent calibrations from Speagle+14. [also Whitaker+14, Schreiber+14].

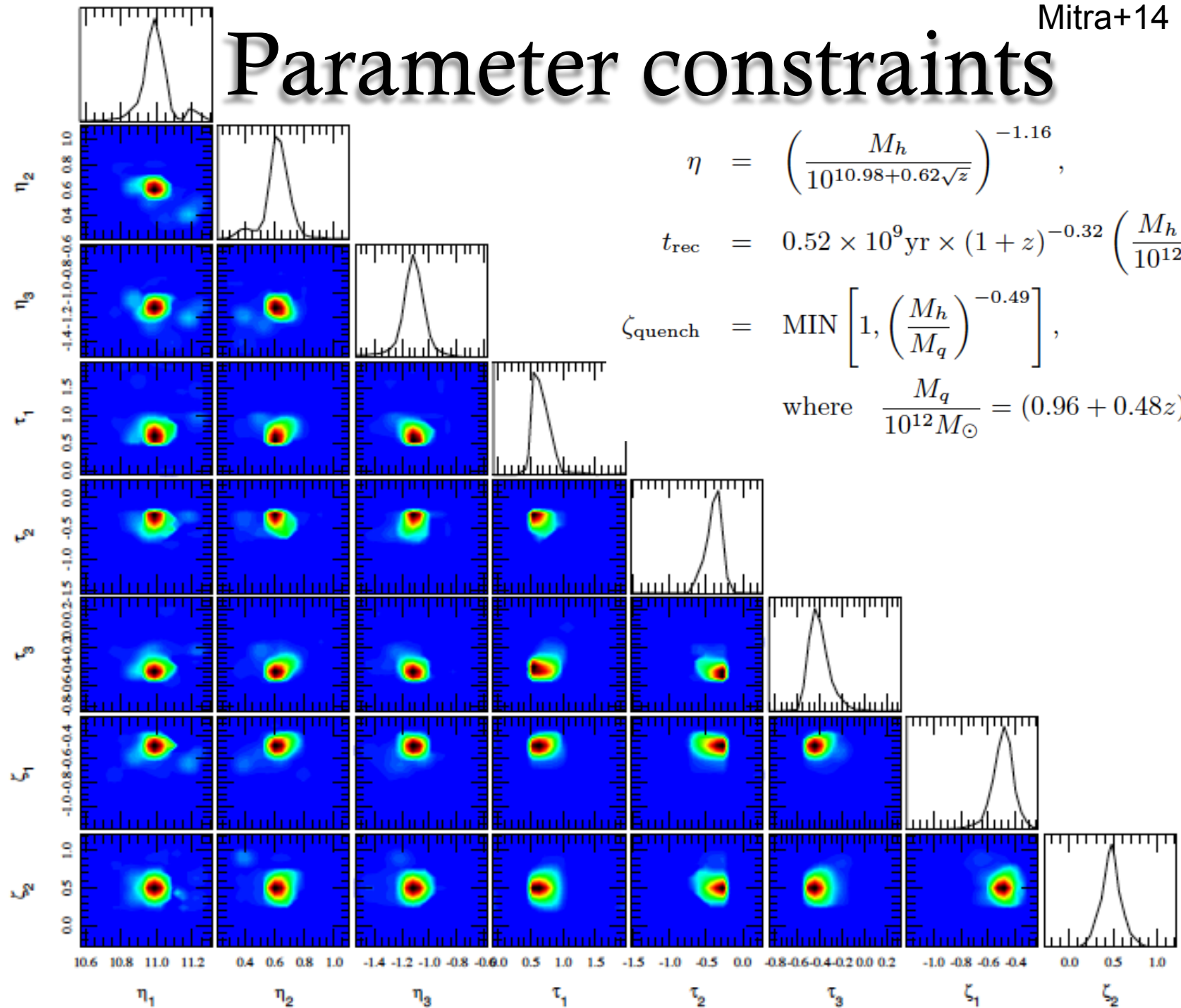
Equilibrium Model: MCMC constraints ($\chi^2 \sim 1.6$)



Fit only to M_* - M_h , SFR- M_* , predict M_* - Z ($\chi^2 \sim 2.0$)

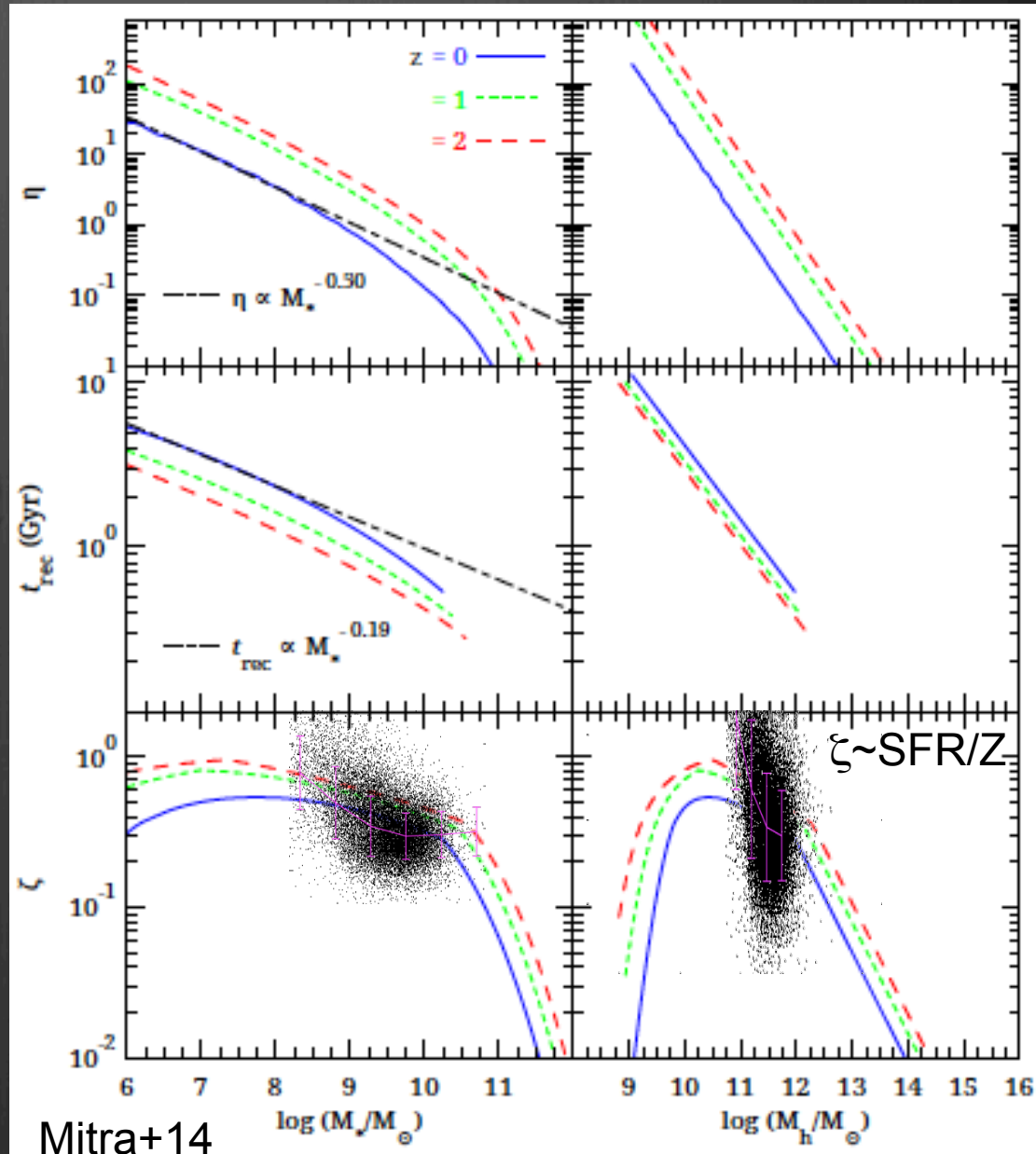


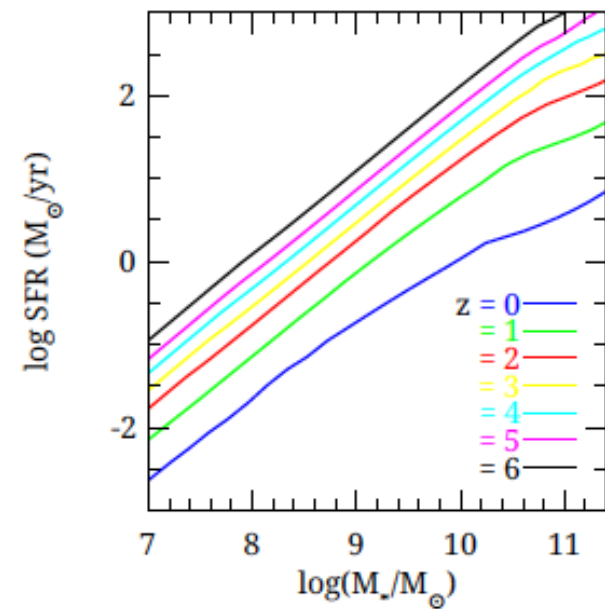
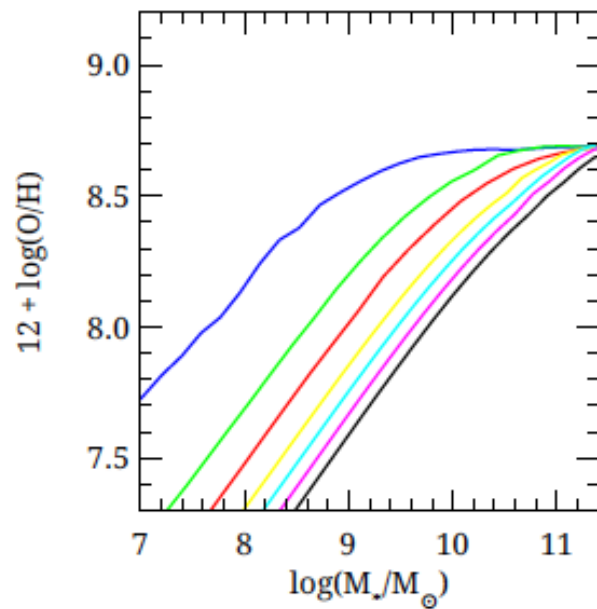
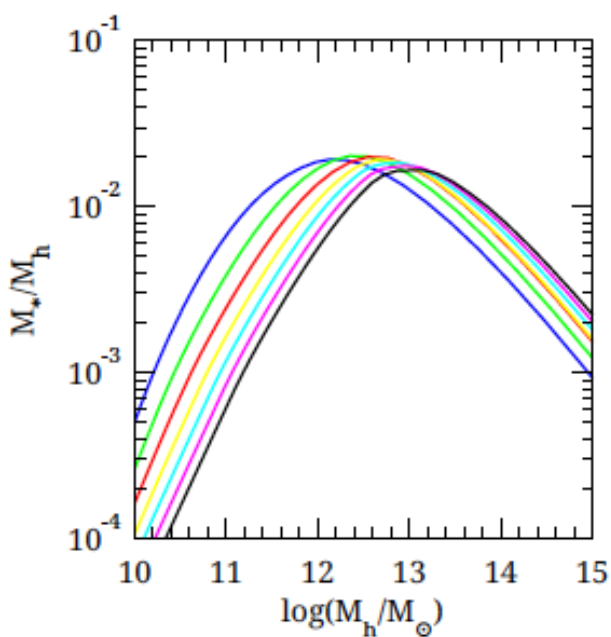
Parameter constraints



Outflows stronger in low-mass galaxies

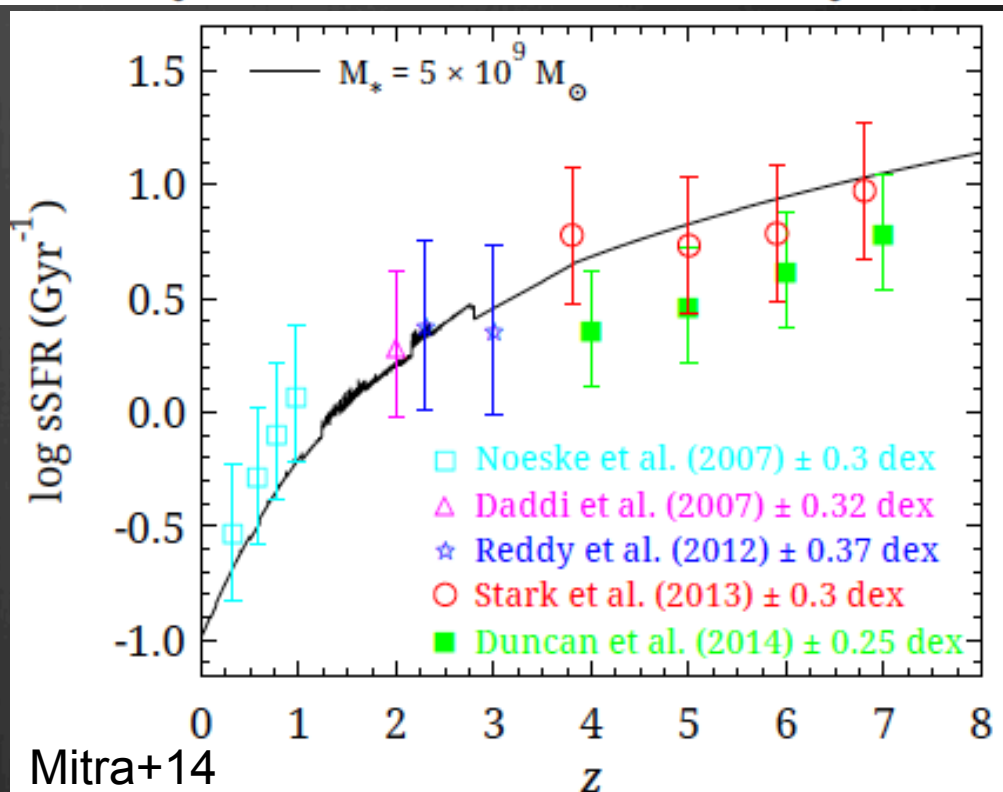
- $\eta \sim M_h^{-1.2} \sim M_*^{-0.3}$
- Stronger at hi-z
- FIRE (Muratov+15):
 $\eta \sim M_*^{-0.35}$,
 amplitude like at $z \sim 2$
- $t_{\text{rec}} \sim M_h^{-0.45} \sim M_*^{-0.2}$
- Opp, RD+08, 10
 best-fit hydro sims:
 $t_{\text{rec}} \sim M_h^{-0.5}$





Evolution of scaling relations

Can now fit sSFR(z) at $z \sim 2$!



Intuition from the Equilibrium Model

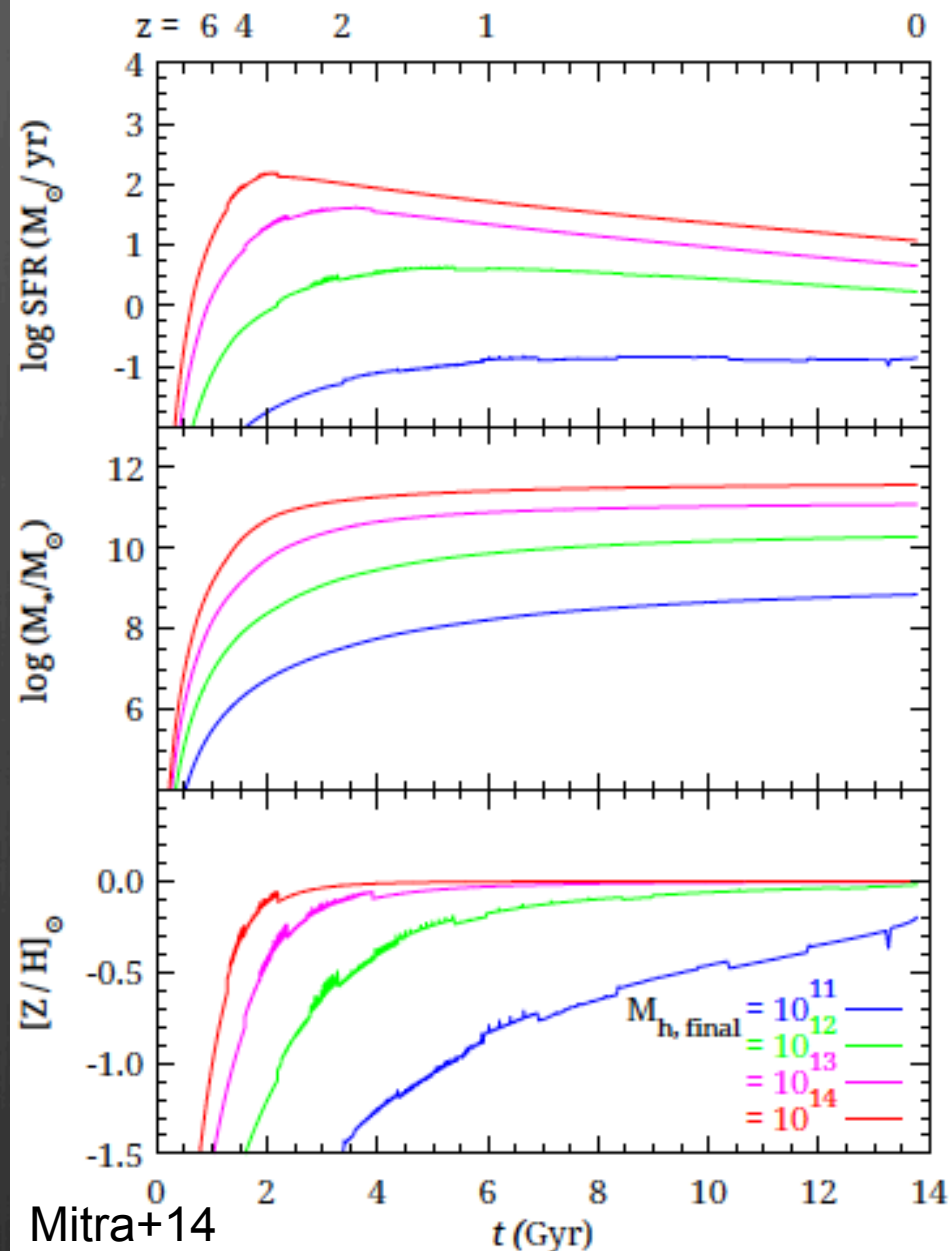
- ⊙ Stellar and metal growth limited by ~~cooling rate and conversion of gas into stars~~
ejective and preventive feedback
- ⊙ Metal & gas content reflects ~~“evolutionary state”~~
gas supply vs. consumption rate
- ⊙ Mergers ~~drive galaxy evolution~~
are subdominant to cold streams for fueling
- ⊙ Galaxies & IGM ~~evolve independently~~
are connected by baryon cycling

Summary

- ⊛ Despite advances, modern SAMs and simulations still have difficulty matching $z \sim 0-2$ data.
- ⊛ The equilibrium model provides a new, simple, and robust approach to phenomenological galaxy formation, based on the simulation-motivated framework of baryon cycling.
- ⊛ This model is able to fit key data across cosmic time with reasonable baryon cycling parameters, in which
 - ⊛ small galaxies eject a larger fraction of accreted gas
 - ⊛ recycling is more rapid in high-mass galaxies
 - ⊛ quenching at a mass that evolves modestly upwards with z

What is it good for?

- Get SFR and Z histories for given M^*, z .
- Serve as input into cosmological sims that include feedback.
- Make predictions for evolution to higher z .
- Populate N-body with “average” galaxies.
- New framework for understanding galaxy formation



Modeling a concordant galaxy population: Status

- ⊙ Simulations, despite many advances, still can't match basic demographics of (M^* , Z , SFR, HI) at $z=0$ let alone at high- z . Improvements still be “trial and error”. *How much farther can we take this?*
- ⊙ SAMs now employ MCMC etc, but even using 20+ parameters can't match all properties. *Is the basic framework flawed?*

Need a simple, robust parameterization that we can easily match to data and that captures insights from simulation.

Equilibrium Relations

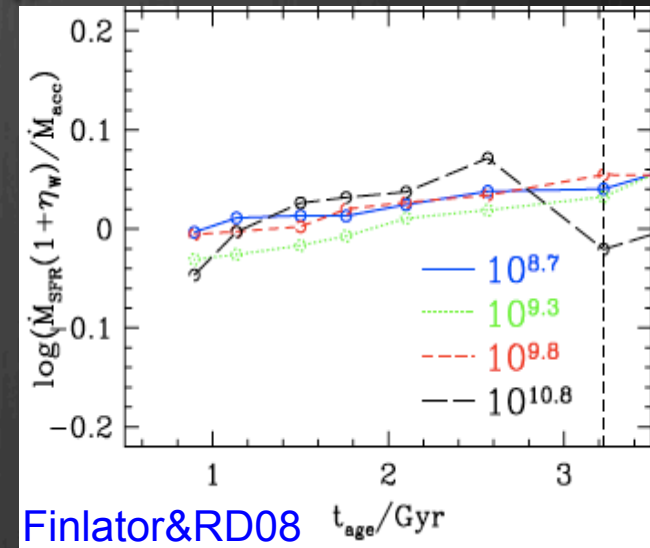
$$\text{Inflow} = \text{SFR} + \text{Outflow} + \frac{d\text{Reservoir}}{dt}$$

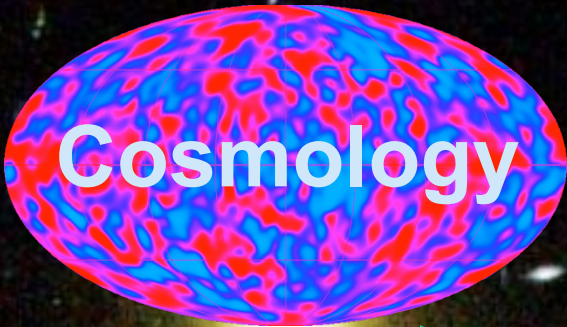
RD+12

- $\text{SFR} = (\zeta \dot{M}_{\text{grav}} + \dot{M}_{\text{recyc}}) / (1 + \eta)$
- $Z = y \text{SFR} / \zeta \dot{M}_{\text{grav}}$
- $f_{\text{gas}} = (1 + (t_{\text{dep}} \text{sSFR})^{-1})^{-1}$

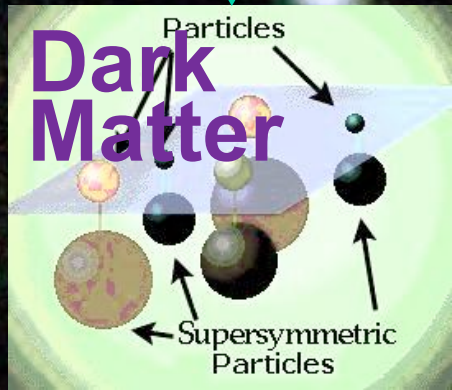
where e.g.

$$\dot{M}_{\text{grav}} \sim f_b M_{\text{halo}}^{1.1} (1 + Z)^{2.25} \quad [\text{Dekel+09}]$$

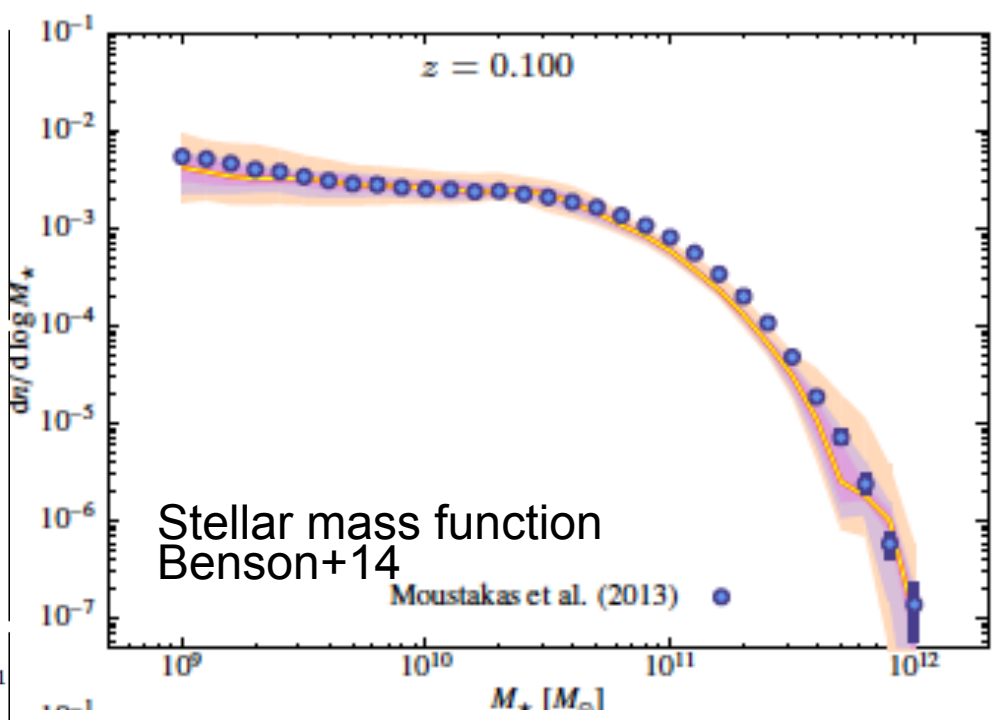
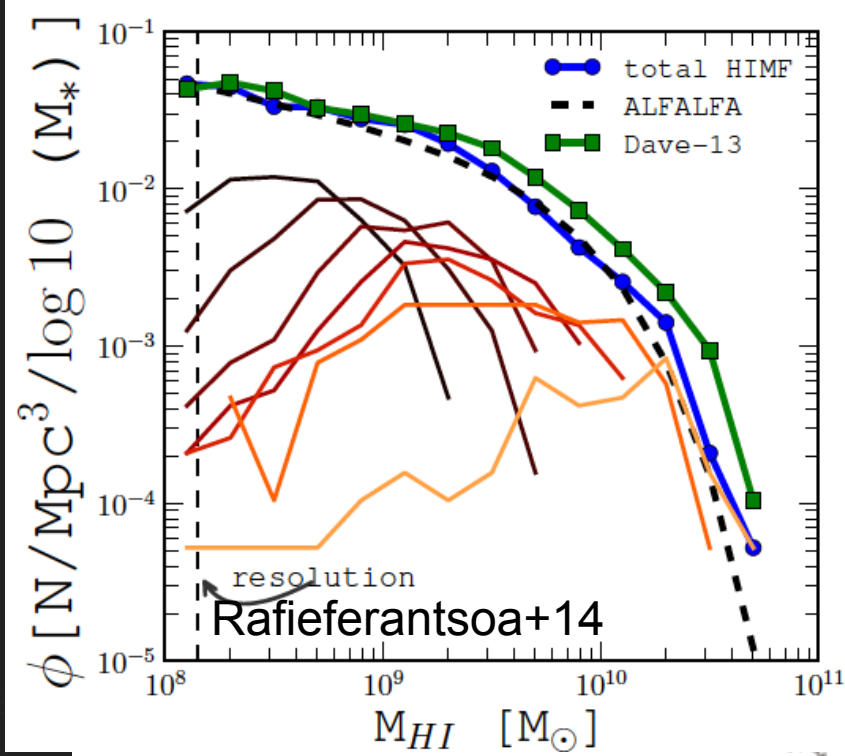




Galaxy Formation

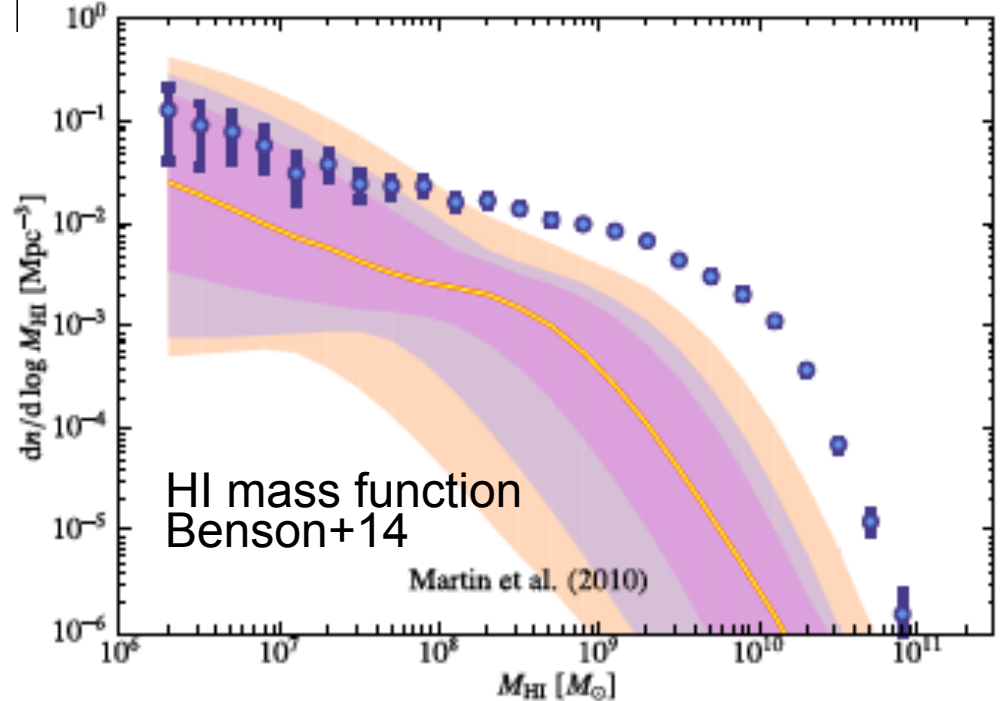


Hubble GOODS field

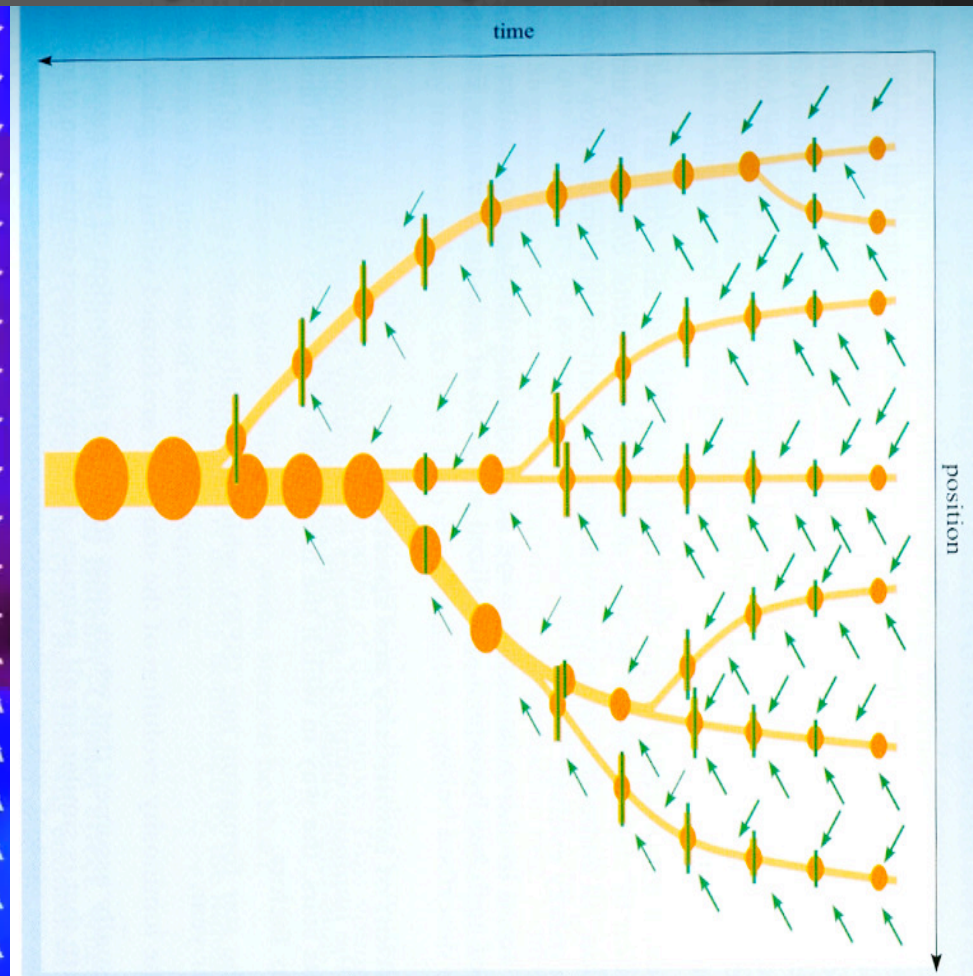
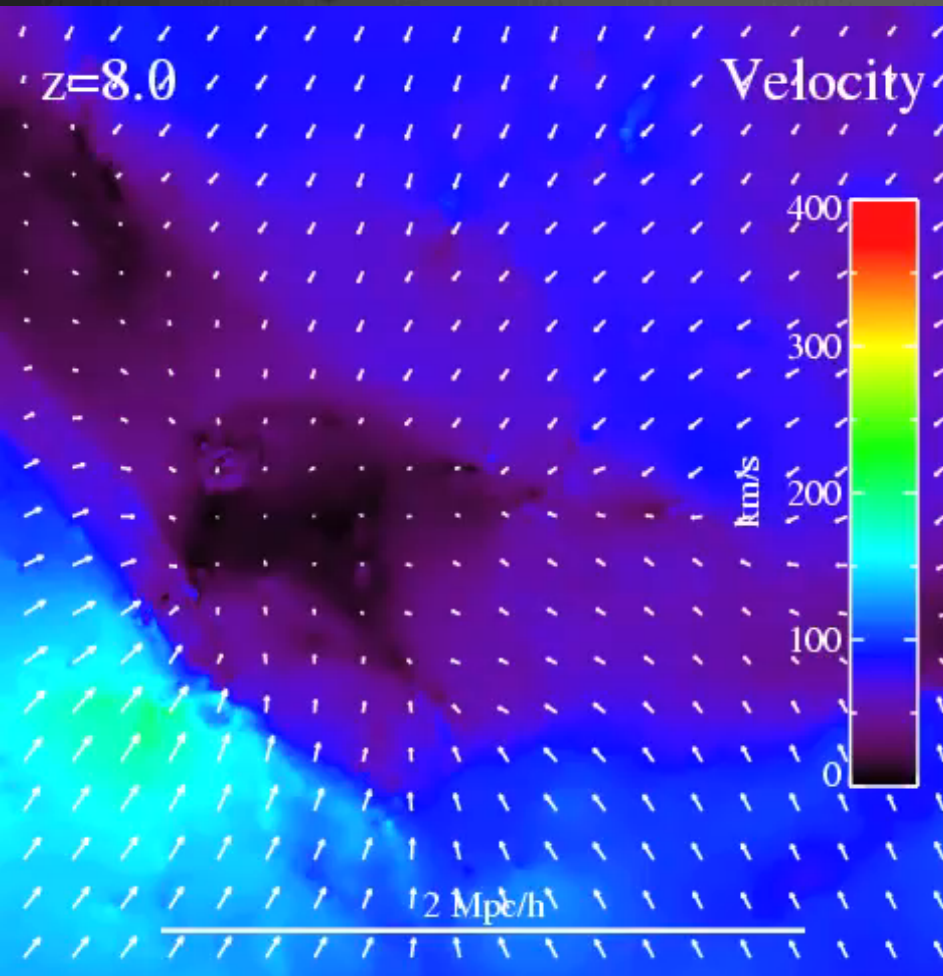


☉ Benson+14: “highly simplified” SAM, “only” 25 parameters. Full MCMC!

☉ Fits GSMF to $z \sim 1$, but badly fails HIMF



Hydro Sims: “Baryon Cycling”



The cycle of inflows, outflows, and re-accretion governs the growth and evolution of galaxies.

What's left to be done?

- ⊛ LOTS!
- ⊛ Consider *scatter*: Fluctuations in gravitational inflow rate, plus(?) excursions via major mergers.
- ⊛ Include *gas* (HI, H₂): Compare/constrain to data
- ⊛ Model *satellites*: Affected by stripping processes.
- ⊛ Examine *environment*: Formation time bias, ...
- ⊛ Compare with hydro simulations w/same parameters
- ⊛ Implement into full N-body simulations.